

SUN/42.5

Starlink Project  
Starlink User Note 42.5

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6 December 1996

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**DAOPHOT — Stellar Photometry  
Package  
Version 1.2  
Introduction**

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## Abstract

DAOPHOT is a stellar photometry package designed by Peter Stetson at DAO to deal with crowded fields. The package performs various tasks including finding objects, aperture photometry, obtaining the point-spread function, and profile-fitting photometry. Profile fitting in crowded regions is performed iteratively, which improves the accuracy of the photometry.

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## 1 Introduction

DAOPHOT is a stellar photometry package designed by Peter Stetson at DAO to deal with crowded fields. The package performs various tasks including finding objects, aperture photometry, obtaining the point-spread function, and profile-fitting photometry. Profile fitting in crowded regions is performed iteratively, which improves the accuracy of the photometry.

This document is an introduction to DAOPHOT II: The Next Generation. It replaces the previous version which is known as DAOPHOT Classic. The main changes concern the choice of the point-spread fitting function and the handling of under-sampled data. This version uses image data written in NDF, the Starlink N-Dimensional Data Format (SUN/33), that contain a DATA\_ARRAY component and have an extension .sdf.

A document written by Peter Stetson (DAOPHOT: A computer program for crowded-field stellar photometry, *Publ.astr.Soc.Pacific*, **99**, 191-222, 1987) gives the background to the algorithms used by DAOPHOT. A user's manual (MUD/9) is also available through Starlink.

DAOPHOT does not directly use an image display, which is one of the reasons why the package has been successfully ported round the world. Three additional routines have therefore been supplied which allow results obtained with DAOPHOT to be displayed on an image device.

The routine DAOGREY (see §3) will display a grey scale image of the data on a suitable device. DAOPLOT (see §4) will indicate the positions of objects found with DAOPHOT on top of the grey image. DAOCURS (see §5) will put up a cursor on the display to allow positions to be measured from the screen. Although KAPPA (SUN/95) routines could be used for the image display and cursor interaction, DAOPHOT does not comply with the Starlink coordinate system convention (SUN/33) and thus any coordinates would be half a pixel out.

DAOPHOT is copyright of Peter Stetson at the Dominion Astrophysical Observatory. The algorithms should not be changed without permission.

## 2 Running DAOPHOT

DAOPHOT is invoked by the simple command:

```
% daophot
```

this should generate a response similar to:

```
DAOPHOT applications are now available -- (Version 1.2)
```

Subsequent use of daophot will place you inside the DAOPHOT executable image, which you can leave with the command:

```
Command: EXIT
```

Inside the executable image you can identify the data to be analysed by DAOPHOT using the command:

```
Command: ATTACH <filename>
```

The file has to be in the NDF format of the Starlink Standard Data Structures (SUN/33). The default file extension is `.sdf`.

At present DAOPHOT can only handle NDF files that have the `DATA_ARRAY` structure at the top level. If the program fails to find the data structure in the container file, then the message:

```
Failed to attach <filename>
```

will be displayed. If you do not expect this message, then this is the moment to seek help from your system manager. If you do not get this message, then you can carry on with DAOPHOT.

The complete set of DAOPHOT commands can be found in the user's manual (MUD/9).

There are three options files which are used by DAOPHOT at various stages of the reduction. There are examples of these in the directory specified by the `DAOPHOT_DIR` environmental variable. The file `daophot.opt` is used when the program starts up. Parameter definitions can be added or removed from the options file as required. An option is specified using a two letter mnemonic followed by its value. If a copy of this file is not present in the default directory, or the parameters have invalid values, then the program will prompt for those parameters that it cannot continue without, namely the read-out noise and the gain. This is signified by the message:

```
Value unacceptable --- please re-enter
```

followed by the parameter prompt. The user has to give values for the read-out noise and gain before the program will continue. Both parameters are used at various stages of processing and the user guide keeps stressing that it is highly advisable to have the most accurate values as possible for these parameters. Although the program will execute with guesstimates for these, the results should be treated with caution. The sections on the **PEAK** and **NSTAR** commands in the user guide suggest a method of checking the validity of the values given.

The file `photo.opt` contains the list of aperture radii used by DAOPHOT when performing aperture photometry.

An alternative to the iterative profile fitting routine **NSTAR** in DAOPHOT has been written by Peter Stetson and is called **ALLSTAR**. It is described in the user manual and can be executed by typing:

```
% allstar
```

The file `allstar.opt` contains parameters for this program.

### 3 DAOGREY – Display an image

DAOPHOT does not directly use an image display, instead it produces image files of intermediate results which can be displayed on any available display system. It is often desirable to indicate the positions of stars on top of the displayed image. To achieve this some special routines have been created.

DAOGREY is the first of these routines and this will display the image data on a suitable device. It is executed by typing:

```
% daogrey
```

The following parameters are requested:

- IN – NDF containing input image** The name of the file containing the image data.
- XSTART – first index of subarray**
- XEND – second index of subarray**
- YSTART – first index of subarray**
- YEND – second index of subarray** These specify the limits in pixel coordinates of the area of the data array to be plotted.  
The choice of values allows subsections of the data to be viewed.
- LOW – lowest data value to plot**
- HIGH – highest data value to plot** These specify the data values corresponding to black and white on the display. The grey levels in between are set automatically.
- DEVICE – Display device** The name of the display device to be used. The image is plotted in the centre of the display, best filling the display area.

### 4 DAOPLOT – Overlay object positions on an image

DAOPLOT will take a file produced by DAOPHOT and overlay the positions of objects in the file on top of the grey image produced by DAOGREY. The positions are indicated by a green cross. The file can be any of the types .coo, .ap, .grp or .nst produced by DAOPHOT. If the file is not suitable then error messages will be given and the routine will end. If the filename was entered wrongly, then just run the task again.

To run DAOPLOT simply type:

```
% daoplot
```

The following parameter is requested:

- PFILE – File with object positions** The name of the file containing the positions of the objects to be indicated. The filename and extension should be given.

## 5 DAOCURS – Overlay a cursor on an image

DAOCURS will display a cursor over the grey image produced by DAOGREY so that positions can be read off the screen. The cursor is positioned with the mouse and the position recorded by pressing any of the keys, except the break key on the mouse. The break key ends the cursor session. Positions are written to the terminal as well as a file.

To run DAOCURS type:

```
% daocurs
```

The following parameter is requested:

**RFILE – File for results** The name of the file to receive the positions selected by the cursor operation.

## 6 Database

DAOGREY, DAOPLOT and DAOCURS have been specially tailored round each other. DAOPLOT and DAOCURS will not at present work on an image displayed with any routine other than DAOGREY. DAOGREY uses the AGI database system (SUN/48) to record the position of the image on the screen. DAOPLOT and DAOCURS use the database file to set up the screen coordinates. When an entire session has been completed the database file can be deleted.